

Color Superconductivity in a Strong Magnetic Field

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SEWM 2006, BNL

Introduction

- The QCD phase diagram at low T and high μ is very rich - many different color superconducting phases, depending on the value of quark masses and chemical potentials (see [Shovkovy's talk](#)).

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Orders of Magnitude

All compact stars support magnetic fields

$B \sim 10^{12} - 10^{14} \text{ G}$ in the surface of pulsars

$B \sim 10^{15} - 10^{16} \text{ G}$ in the surface of magnetars

There is an upper limit to the star magnetic field
(compare gravitational and magnetic energies)

$$B_{\text{max}} \sim 1.4 \times 10^{18} \left(\frac{M}{M_{\odot}} \right) \left(\frac{10 \text{ km}}{R} \right)^2 \text{ G}$$

Color Flavor Locking Phase

For $m_q \approx 0$ ($m_s < 2\sqrt{\mu\Delta}$).

$$\langle q_L^{ia} q_L^{jb} \rangle = \Delta_A \epsilon^{ijk} \epsilon_{abk}$$

$a, b = 1, 2, 3$ flavor indices $i, j = 1, 2, 3$ color indices
(Alford, Rajagopal, Wilczek, '98)

Local and global symmetries are spontaneously broken

$$SU(3)_C \times SU(3)_L \times SU(3)_R \times U(1)_B \rightarrow SU(3)_{C+L+R}$$

→ several similitudes with vacuum QCD

- There are Goldstone bosons associated to SSB of chiral symmetry ($\pi^0, \pi^\pm, K^0, \bar{K}^0, K^\pm, \eta$)
- One Goldstone boson for the breaking of $U(1)_B$

CFL matter in a external magnetic field

Is a CFL color superconductor also an electromagnetic superconductor ?

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The condensates $\langle qq \rangle$ also break spontaneously $U(1)_{\text{e.m.}}$.

But there is a combination of electromagnetism and a $U(1)$ subgroup of $SU(3)$ that remains unbroken

- Seven gluons and one combination of gluon and photon are massive.
- One combination of gluon and photon is massless

$$\tilde{G}_\mu^8 = \cos \theta_{\text{CFL}} G_\mu^8 + \sin \theta_{\text{CFL}} A_\mu ,$$

$$\tilde{A}_\mu = -\sin \theta_{\text{CFL}} G_\mu^8 + \cos \theta_{\text{CFL}} A_\mu .$$

$$\cos \theta_{\text{CFL}} = \frac{\sqrt{3}g}{\sqrt{3g^2 + 4e^2}}$$

CFL and the in-medium electromagnetism

All quark, gluon and meson charges are integral

$$\tilde{e} = e \cos \theta_{CFL}$$

| s_1 | s_2 | s_3 | d_1 | d_2 | d_3 | u_1 | u_2 | u_3 |
|-------|-------|-------|-------|-------|-------|-------|-------|-------|
| 0 | 0 | - | 0 | 0 | - | + | + | 0 |

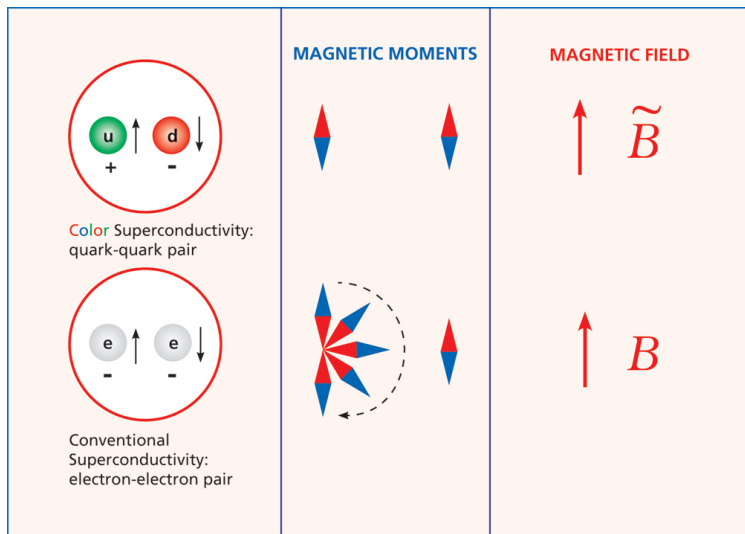
the “rotated” photon is massless; but the medium still modifies its propagation properties

$$\tilde{v} = 1/\sqrt{\tilde{\epsilon}} < 1 \quad \text{Litim and C.M. 01}$$

the CFL medium acts as a transparent insulator, with interesting reflexion/refraction properties

C.M. and Rajagopal, 01

Influence of a Strong Magnetic Field in Superconductivity



Influence of B in fermion pairing

- In an electromagnetic superconductor
a strong B field tends to break the condensate
- In the CFL color superconductors
the penetrating field tends to stabilize the condensate.
- Magnetic catalysis of a chiral (fermion-antifermion)
condensate at zero density , even at weak coupling
Gusynin, Miransky and Shovkovy, 94,
 \Rightarrow dimensional reduction of the pairing dynamics at the
lowest Landau level

MCFL

Writing only the antisymmetric gaps

$$\langle q_L^{ia} q_L^{jb} \rangle = \Delta_A \epsilon^{ij3} \epsilon_{ab3} + \Delta_A^B (\epsilon^{ij1} \epsilon_{ab1} + \epsilon^{ij2} \epsilon_{ab2})$$

Symmetry breaking pattern

$$\begin{aligned} SU(3)_C \times SU(2)_L \times SU(2)_R \times U^{(-)}(1)_A \times U(1)_B \times U(1)_{\text{e.m.}} \\ \rightarrow SU(2)_{C+L+R} \times \tilde{U}(1)_{\text{e.m.}} \end{aligned}$$

→ several similitudes with vacuum QCD in an external B

Miransky and Shovkovy (2002)

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- Use a Nambu-Jona-Lasinio (NJL) model, inspired by one-gluon exchange, to study more moderate densities (Λ UV cutoff)

$$\mathcal{L}_I = \frac{g^2}{\Lambda^2} \bar{\psi} \gamma^\mu \lambda^A \psi \bar{\psi} \gamma_\mu \lambda^A \psi$$

$$\mu \sim 400 - 500 \text{ MeV} , \quad \Delta \sim 10 - 50 \text{ MeV}$$

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- Interesting gluon dynamics at weak coupling and moderate fields see **Ferrer and Incera 06**

Some few technical details: Ritus method

$$\Pi_{\mu}^{(\pm)} = i\partial_{\mu} \pm \tilde{e}\tilde{A}_{\mu}$$

$$(\Pi^{(\pm)} \cdot \gamma) E_q^{(\pm)}(x) = E_q^{(\pm)}(x) (\gamma \cdot \bar{p}^{(\pm)})$$

$$\bar{p}^{(\pm)} = (p_0, 0, \pm \sqrt{2|\tilde{e}\tilde{B}|k}, p_3) \quad k \text{ labels the Landau levels}$$

$$E_q^{(\pm)}(x) = \sum_{\sigma} E_{q\sigma}^{(\pm)}(x) \Delta(\sigma)$$

$$\Delta(\sigma) = \text{diag}(\delta_{\sigma 1}, \delta_{\sigma -1}, \delta_{\sigma 1}, \delta_{\sigma -1}), \quad \sigma = \pm 1$$

$$E_{p\sigma}^{(\pm)}(x) = \mathcal{N}_{n_{(\pm)}} e^{-i(p_0 x^0 + p_2 x^2 + p_3 x^3)} D_{n_{(\pm)}}(\varrho_{(\pm)}) ,$$

$D_{n_{(\pm)}}(\varrho_{(\pm)})$: parabolic cylinder functions

$$\varrho_{(\pm)} = \sqrt{2|\tilde{e}\tilde{B}|} (x_1 \pm p_2/\tilde{e}\tilde{B}) ,$$

MCFL

We have solved the gap equations in an effective NJL model, inspired by one-gluon exchange, and for strong magnetic fields $\tilde{e}\tilde{B} > \mu^2/2$

(then all the charged quarks are in the lowest Landau level)

$$\Delta_A^B \sim 2\mu \exp\left(-\frac{3\Lambda^2\pi^2}{g^2(\mu^2 + \tilde{e}\tilde{B})}\right)$$

and $\Delta_A \ll \Delta_A^B$

to be compared with the CFL fermionic gap

$$\Delta_A^{\text{CFL}} \sim 2\sqrt{\delta\mu} \exp\left(-\frac{3\Lambda^2\pi^2}{2g^2\mu^2}\right)$$

Magnetic catalysis of the diquark condensate

BCS behaviour of the gap

$$\Delta \propto \exp(-1/G^2\rho)$$

ρ : density of states close to the Fermi surface

a strong magnetic field increases the density of states of the charged quarks close to the Fermi surface!

from $\mu^2/2\pi^2 \Rightarrow \tilde{e}\tilde{B}/2\pi^2$.

orders of magnitude for the effect to be relevant

$$\tilde{e}\tilde{B} \sim 10^{18} G$$

for astrophysical applications one should look to more moderate fields - solving the gap equations for moderate fields requires a numerical analysis, with the inclusion of higher Landau levels

MCFL Low energy effective field theory

$$\Sigma = XY^\dagger = \exp\left(i\frac{\Phi}{f_{\pi,B}} + i\phi_0\right), \quad \Phi = \phi_A \sigma^A, \quad A = 1, 2, 3$$

The external magnetic field introduces a strong anisotropy in the system

$$\mathcal{L} = \frac{f_{\pi,B}^2}{4} \left(\text{Tr} \left(\partial_0 \Sigma \partial_0 \Sigma^\dagger \right) + \left(v_\perp^2 g_\perp^{ij} + v_\parallel^2 g_\parallel^{ij} \right) \text{Tr} \left(\partial_i \Sigma \partial_j \Sigma^\dagger \right) \right)$$

The parameters of the low energy effective theory have to be computed!

Estimates

For which values of the magnetic field the superconductor is in CFL or MCFL phases?

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Estimate based on the low energy effective field theory C.M.,05

CFL: 9 Goldstone bosons \rightarrow MCFL: 5 Goldstone bosons

the magnetic field makes the charged GB π^\pm, K^\pm **massive**

$$M_{\pi^\pm}^2 = M_{K^\pm}^2 \propto \frac{(\tilde{e}\tilde{B}^{\text{ext}})^2}{f_\pi^2}$$

when the mass of those are of order 2Δ they decouple

$$\tilde{e}\tilde{B} \sim 2f_\pi\Delta, \quad \tilde{e}\tilde{B} \sim 10^{16} \text{ G}$$

Conclusions

- An applied external strong magnetic field leads to a new color superconducting phase. B has both qualitative and quantitative effects on color superconductivity!
The low energy properties, including transport properties, interactions with neutrinos, cooling, etc, will differ.
- The dynamics of the magnetic field will also be so different.